Potential and limitations of CsBi₃I₁₀ as photovoltaic material

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Herein we demonstrate the dry synthesis of CsBi₃I₁₀ both as freestanding material as well as in the form of homogeneous thin films, deposited by thermal vacuum deposition. Chemical and optical characterization shows high thermal stability, phase purity, and photoluminescence centered at 700 nm, corresponding to a bandgap of 1.77 eV. These characteristics make CsBi₃I₁₀ a promising low-toxicity material for wide bandgap photovoltaics. Nevertheless, the performance of this material as a semiconductor in solar cells remains rather limited, which can be at least partially ascribed to a low charge carrier mobility, as determined from pulsed-radiolysis, time-resolved microwave conductivity. Further developments should focus on understanding and overcoming the current limitations in charge mobility, possibly by compositional tuning through doping and/or alloying, as well as optimizing thin film morphology which may be another limiting factor.

1. Introduction

Lead halide perovskites (LHPs) have attracted a lot of attention over the past decade as a new family of semiconducting materials. Single-junction solar cells made thereof have been demonstrated with power conversion efficiencies (PCEs) of 25.2%, very close to record values for silicon photovoltaics (PV).¹ Furthermore, LHPs possess a number of advantages over silicon and other inorganic semiconductors, such as a high degree of bandgap tunability and low-temperature processing. On the other side, LHPs, as well as other PV materials such as CdTe or GaAs, are composed of highly-toxic elements (Pb, Cd, As), which might undermine their widespread application.² Therefore, in parallel with the development of high-efficiency LHP solar cells, the focus is now shifting towards finding and developing low-toxicity, alternative semiconductors. While the fast pace at which these are proposed is exciting, it is also important to assess their potential critically.³ In this context, a few publications have suggested CsBi₃I₁₀ as a promising candidate, reporting PCEs from 0.13% to 1.5%.^{4–7}

Motivated by these results, we demonstrate hereafter that CsBi₃I₁₀ can be synthesized as a free-standing powder (nontemplated synthesis) as well as thin films by fully-dry approaches, thus also avoiding the toxicity of most common organic solvents. $^{8\text{--}13}$ CsBi_3I_{10} powders are synthesized by a simple mechanochemical route (neat ball-milling) whereas thin films are prepared by dual-source thermal vacuum codeposition of CsI and Bil₃. Physical and chemical characterizations reveal that the as-prepared materials show crystallinity and stoichiometry corresponding to the expected CsBi₃I₁₀ phase, with an optical bandgap of approximately 1.77 eV. Semiconductors with such a bandgap are highly sought after for front-cell absorbers in tandem solar cells.^{14–16} The thin films are also found to be rather smooth (RMS roughness below 7 nm) with homogeneous photoluminescence as evidenced by hyperspectral imaging.

Despite these encouraging characteristics, the implementation of this material in solar cells results in rather low photovoltaic performance, in line with previous reports (PCE = 0.13%).^{6,17} Importantly, we find that with the chosen electron- and hole-transport layers, the direct *p-n* junction (without CsBi₃I₁₀) yields a higher PCE. These low performances may be partially ascribed to film morphology (small grains). Nevertheless, pulsed-radiolysis time-resolved microwave conductivity (PR-TRMC) measurements also reveal that despite an extraordinary long charge carrier lifetime of up to tens of microseconds (pointing to a low defect density), the effective mobility is limited to about 0.024 cm²/Vs, two orders of magnitude lower than that of LHPs.¹⁸ In view of this, it is worth reconsidering the potential of this material. Indeed, while the direct bandgap of 1.77 eV (photoluminescence centered at 700 nm) and the possibility of the solvent-free synthesis demonstrated herein are promising, the intrinsic charge carrier

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mobility should be greatly enhanced (possibly through doping or other compositional tuning) for an efficient use of this material in photovoltaics.

2. Results and discussion

Mechanochemical synthesis of free-standing CsBi₃I₁₀

Non-templated CsBi₃I₁₀ powders were synthesized by neat ballmilling of CsI and BiI₃ in 1:3 molar ratio under inert atmosphere (see Experimental Section for details). X-ray diffraction, optical characterization, and thermogravimetric analysis of the asobtained materials are presented in Figure 1.



Fig. 1 (a) X-ray diffractogram of the $\rm CsBi_{3}I_{10}$ powders (green line) and reference patterns for CsI (red; ICSD 44938) and BiI₃ (blue; ICSD 56573) (b) normalized absorbance and photoluminescence spectra (excitation wavelength 522 nm) and (c) thermogravimetric analysis (TGA) of CsBi_{3}I_{10} powders.

The X-ray diffractogram (XRD) of the powder after ballmilling is shown in Figure 1a. The first observation that can be made is that several major peaks of the binary precursors are absent. Since the powder sample is not expected to have any major preferential orientation which could otherwise explain the suppression of selected peaks, it is reasonable to conclude that CsI and Bil₃ have fully reacted. While we did not find a reference pattern for $CsBi_3I_{10}$ in the Inorganic Crystal Structure Database (ICSD) or other databases, we note that the diffractogram obtained here matches the expected peaks and relative intensities for $CsBi_3I_{10}$, based on previously published data.⁶ In this article, Johansson et al. suggest that $CsBi_3I_{10}$ crystallizes in a structure that is closely related to that of Bil_3 and Cs₃Bi₂I₉, which is another stable ternary phase. This explains some similarities in the XRD patterns between the three phases. Such similarities have also been observed by others.4,5 Nevertheless, if we compare our diffractogram to the reference pattern of Cs₃Bi₂I₉ (Figure S1) it becomes apparent that several major peaks of the latter are not present, pointing to the absence of $\mathsf{Cs}_3\mathsf{Bi}_2\mathsf{I}_9.$ We can thus conclude that $\mathsf{CsBi}_3\mathsf{I}_{10}$ is formed, with the full reaction of binary precursors and without the formation of other ternary phases.

This conclusion is further supported by optical characterization. The optical absorption and photoluminescence (PL) spectra are presented in Figure 1b. We observe an absorption onset and PL peak centered around 700 nm, corresponding to an optical bandgap of 1.77 eV (see Tauc plot in Figure S2). This is in good agreement with previously-reported values^{17,19} and different from the emission of Cs₃Bi₂I₉ which is centered at 650 nm.²⁰

Importantly, the difference between the cut-off of the absorption spectrum and the PL maximum (Stokes-shift) is small. This feature, together with the steep absorption onset, is expected for direct bandgap semiconductors, where the charge carrier generation and recombination occur at the band edges. The direct bandgap of 1.77 eV for as-prepared CsBi₃I₁₀ is of particular interest in wide bandgap solar cells for tandem applications.²¹

In order to assess the thermal stability of this material, we have also conducted thermogravimetric analyses (TGA; Figure 1c). The curve shows a first weight loss of 68% around 275 °C, followed by a second loss from the remaining 32% to 13% of the initial mass, around 475 °C. These losses can be rationalized by considering the following degradation mechanism:

- (1) 3 CsBi₃I₁₀ (s) = Cs₃Bi₂I₉ (s) + 7 BiI₃ (g), around 275 °C
- (2) Cs₃Bi₂I₉ (s) = 3 CsI (s) + 2 BiI₃ (g), around 475°C

Considering the molecular weights and stoichiometry involved in each of the two steps, the observed weight losses are fully consistent with the expected values. The thermal stability of $CsBi_3I_{10}$ up to around 250 °C is significantly higher than that of the most common hybrid organic-inorganic lead halide perovskites.^{22,23} Moreover, it is well beyond the usual range of operation of most optoelectronic devices, which ensures that no thermal-induced degradation should occur.

Vacuum-deposited thin films

In order to fabricate PV devices, we tested the deposition of CsBi₃I₁₀ as a thin film. Reported procedures so far are based on the dissolution of the inorganic salts in an organic solvent and subsequent spin-coating.²⁴ Nevertheless, typical solvents such as DMF used for this approach are highly toxic and generate a considerable amount of waste. Hence, using them partly defeats the purpose of developing this low-toxicity material. Therefore, we used thermal vacuum deposition instead. Ideally, as-synthesized CsBi₃I₁₀ powders could be used to fabricate thin films by single-source vacuum deposition (SSVD).25,26 Nevertheless, our attempts at depositing CsBi₃I₁₀ films by SSVD resulted in the decomposition of $CsBi_3I_{10}$ into $Cs_3Bi_2I_9$ + BiI_3 , with the latter subliming at much lower temperature and with a higher deposition rate than the former. Consequently, thin films were fundamentally composed of Bil₃, while Cs₃Bi₂I₉ was found to remain in the crucible (see Figures S3-S4). These observations are consistent with the first degradation mechanism observed by TGA (see discussion above) and highlight the fact that it cannot be avoided here despite fast heating in vacuum.

Therefore, we used an alternative method to prepare thin films by co-deposition of CsI and Bil₃ from two thermal sources in a vacuum chamber (base pressure 10^{-6} mbar), each set at a suitable temperature to adjust the deposition rates (see Experimental Section for details).

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In situ, temperature-dependent XRD (Figure 2a) shows the evolution of the CsBi₃I₁₀ thin-film structure from room temperature (RT) to 150 °C, which is different from the diffractogram of the same compound in the powder state (Figure 1a). Nevertheless, a comparison with published data from Johansson et al.⁶ shows that the main peaks observed here around $2\theta = 12.7^{\circ}$ and $2\theta = 25.6^{\circ}$ may be indexed as the (003) and (006) planes of the crystal structure suggested therein. This reveals that the thin film has a preferential orientation with the [001] direction being perpendicular to the substrate. Such a



Fig. 2 Characterization of co-evaporated CsBi₃1₁₀ thin film (the film is previously annealed at 80 °C for b, c, and d; final thickness of 80 nm). (a) XRD at variable temperature, (b) XPS survey spectrum, (c) normalized absorbance and photoluminescence spectra (excitation wavelength 532 nm), (d) scanning electron microscope (SEM) image and optical photograph (inset).

texture (preferential orientation) is common in thin films and explains the differences observed between the XRD signals of powders (Figure 1a) and thin films (Figure 2a) of the same material.27 Interestingly, the diffractogram remains mostly unchanged from room temperature to 150 °C, denoting that no phase transformation occurs in this range. Peak shifts are also minimal (around 0.1° - $0.2^\circ),$ pointing to a very limited thermal expansion of this material. It must be noted, though, that other chemical variations such as the presence (before or after annealing) of metallic bismuth, as observed by others,⁴ could not be discarded. In order to determine the ratio of the different elements present in the thin film, X-ray photoelectron spectroscopy (XPS) was performed. Films were previously annealed at different temperatures to check if this process can help the formation of the desired stoichiometry. The different ratios of Cs, Bi and I calculated for each temperature tested (pristine, 50 °C and 80 °C) are reported in Table S1. The ratios might suggest that an increase in the annealing temperature of the films favors the formation of a stoichiometry closer to the nominal one, as the Cs:Bi:I ratio obtained at 80 °C (Figure 2b) is 1.0: 3.4: 10. However, it must be noted that XPS is a surfacesensitive technique (depth of analysis around 2 nm - 10 nm). Hence, the different values obtained in this temperature range are most likely due to the diffusion of different species to the



Fig. 3 (a) PL spectra (averaged, bright and dark parts) of the surface of the coevaporated thin-film, represented in (b). The blue dot corresponds to the dark spot and the red one, to the bright one. NB: Green spectrum in panel (a) is the same as plotted in Figure 2c. Excitation wavelength: 532 nm.

surface. Optical absorption and photoluminescence (Figure 2c) spectra show the expected features for $CsBi_3I_{10}$, which are significantly different from those of $Cs_3Bi_2I_9$, as previously discussed. A close look at absorption and emission properties of the thin film and powder sample show slight differences (see Figures S2 and S5). Indeed, the absorption onset (or onset of Tauc plot) is blue-shifted by about 70 meV (Figure S2) in the film. Similar thickness-dependent blue-shifts have previously been observed on thin films in this range of thicknesses.^{28,29} This behavior has been ascribed to the small grain size present in thinner films. On the contrary, the powder is expected to be highly polydisperse in crystal size. This is coherent with a red-shifted absorption dominated by the larger crystallites and a slightly wider photoluminescence peak (see Figure S5).

Concerning the film morphology, from a macroscopic point of view, a smooth surface is observed without any obvious inhomogeneity over a 3 cm x 3 cm substrate (see Figure 2d, inset). Scanning electron microscopy (SEM; Figure 2d) reveals a small-grained surface (typical grain size below 100 nm), again without any obvious inhomogeneity or pinholes at the microscale. Furthermore, the surface roughness was evaluated by atomic force microscopy (AFM; Figure S6) and a root mean squared value (RMS) of 6.95 nm was determined over an area of 2.6 µm². The homogeneity of the thin film was further evaluated in terms of optical characteristics through hyperspectral photoluminescence microscopy (Figure 3). The PL spectra for two different regions of the film (marked as blue and red) are presented along with the average PL spectra for the full field of view. The film presents regions with higher and lower PL intensity as directly observed from the brighter and darker spots in Figure 3b and the corresponding different intensities of the PL spectra (Figure 3a). Nevertheless, the differences in intensity are less than one order of magnitude. Furthermore, the PL peak maximum and width are very similar. Again, this points to a high homogeneity in the vacuum-deposited thin film.

In order to gain more insight into the energy levels of this material, air photoemission spectroscopy (APS) was measured (see Figure S7). The ionization energy, which corresponds to the valence band maximum (VBM) for an intrinsic semiconductor such as $CsBi_{3}I_{10}$, is found to be -6.0 eV, in close agreement with previously reported values for these species (-5.9 eV) and measured with different techniques.^{19,30} The energy

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corresponding to the conduction band maximum (CBM) can be estimated to -4.2 eV by subtracting the bandgap energy to the VBM, as shown in Table S2.

Implementation in solar cells

Given the promising characteristics of co-evaporated CsBi₃I₁₀ thin films for photovoltaics, we implemented them in singlejunction solar cells. Indium tin oxide (ITO)-coated glass substrates were used as the transparent electrode. A suspension of SnO₂ nanoparticles was deposited by spin-coating onto ITO (obtaining a 30 nm thick film) and then covered with a layer of C₆₀ (10 nm) by thermal evaporation. Then, a 210 nmthick film of CsBi₃I₁₀ was deposited by co-evaporation and annealed at different temperatures. Subsequently, N4,N4,N4',N4'-tetra([1,1'-biphenyl]-4-yl)-[1,1':4',1'-terphenyl]-4,4'-diamine (TaTm, 15 nm) and MoO₃ (5 nm) were deposited on top and the device was finished with a gold electrode (100 nm). Current density-voltage (J-V) curves were collected under illumination from a solar simulator (Figure 4a). Independently on the thermal annealing of the $CsBi_{3}I_{10}$ film, the devices performed poorly with a PCE around 0.12%-0.13%. Surprisingly, the equivalent device without CsBi₃I₁₀ (that is with direct contact between C₆₀ and TaTm) outperformed these. These poor performances might be partly ascribed to the smallgrained morphology of the CsBi₃I₁₀ films as seen by SEM (Figure 2d). Indeed, the moderate gain in short-circuit current after annealing could indicate a beneficial effect from grain sintering. This can be observed by the sharpening of X-ray diffraction peaks, denoting larger crystallites (see Figure S8). Nevertheless,

performances remain poor in all annealing conditions. Furthermore, we exclude misalignment of the energy levels of the transport layers with CsBi₃I₁₀ as responsible for the large performance losses observed here, as depicted in Figure S9. In order to shed more light into the poor performance of CsBi₃I₁₀based devices, we conducted pulse-radiolysis time-resolved microwave conductivity (TRMC) measurements (see Figure 4b, c). Through these measurements, it was determined that the mobility of charge carriers in the material is around 0.024 cm²/Vs. This value is about two to three orders of magnitude lower than common values reported for lead halide perovskites, especially measured with the same technique, 18,31-33 and may explain the low short-circuit current obtained in the solar cells. As the exact crystal structure could not be solved, it is not straightforward to give a precise reason for this low charge mobility. Nevertheless, the lifetime of the charge carriers is extraordinarily long (up to tens of microseconds as evidenced in Figure 4c). This long lifetime suggests that the density of defects in the material is low. This is supported by the pulse length dependence (variation of the initial concentration of charge carriers) experiments that show that the recombination dynamics follow second-order recombination (the height of the conductivity signal and lifetime decay is faster as more charge carriers are generated in the material) without indication of trap filling during the pulse. This latter behavior has been previously observed in lead halide perovskites studied with the same experimental technique.^{32,33}



Fig. 4 (a) J–V characteristics under 1 sun illumination of different solar cells made with CsBi₃I₁₀ annealed at different temperatures as well as without CsBi₃I₁₀. (b and c) Pulse radiolysis TRMC of CsBi₃I₁₀ as a function of pulse length (which results in initial concentrations of charge carriers from 1x10¹⁶ cm⁻³ to 2x10¹⁷ cm⁻³). Panel (a) and (b) vary by time scale only.

Therefore, it seems that the low mobility should be addressed rather by doping or another compositional tuning than by passivating quenching traps, which do not seem to be prevalent.

Conclusions

In summary, we have proposed an alternative, faster and greener method for the synthesis of $CsBi_3I_{10}$, a low-toxicity ternary metal halide with potential use in optoelectronics. This material shows high thermal stability in air and

photoluminescence centered at 700 nm, which endorses its application in the fabrication of photovoltaic devices. In addition, we have achieved thin-film deposition of the same material through thermal vacuum co-deposition and implemented it in solar cells. Nevertheless, these devices performed poorly, probably as a result of the low charge carrier mobility demonstrated by pulsed-radiolysis microwave conductivity measurements. Ongoing developments are focused on a precise understanding of the causes of this low mobility and possible solutions, including compositional

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engineering in multi-cation and multi-anion derivatives as well as doping.

Conflicts of interest

There are no conflicts to declare.

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Notes and references

‡ Electronic supplementary information (ESI) available: synthesis and characterization; thin-film deposition; device fabrication; TMRC; SSVD sub-products and XRD identification; XPS analysis of co-deposited thin-films; air photoemission spectroscopy.

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